Statistical thermodynamics for MD and MC simulations

knowing 2 atoms and wishing to know 10^{23} of them

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Thermodynamic properties of molecular systems

- how is the thermodynamic equilibrium characterized?
- which quantities are of interest?

statistical mechanics: the way from the properties of particles to the thermodynamic properties of ensembles via the partition function

- how to derive the ensemble partition function from the partition function of a single molecule?
- how is partition function connected to phase-space density?
- MD simulation provides an alternative way to thermodynamic quantities
- it is difficult to obtain free energies from normal simulations

Equilibrium and spontaneous processes

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classical thermodynamics → thermodynamic equilibrium and spontaneous process says which quantites are maximized/minimized in equilibrium and show a definite change during spontaneous processes
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microcanonical ensemble:

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equilibrium reached if entropy S is maximized process is spontaneous if entropy increases: \Delta S > 0
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canonical ensemble – more complex because we need to consider system of interest together with surroundings (to identify equilibrium and spontaneity) calculation on the supersystem impossible – an alternative needed

Free energy / enthalpy – fundamental property

How to keep our molecular system and drop the surroundings? introduce a new thermodynamic function:

- Helmholtz free energy F in canonical NVT ensemble
- Gibbs free energy/enthalpy *G* in NPT ensemble

$$F = U - TS$$
 $G = H - TS = U + pV - TS$

- F = F(T, V) or G = G(T, P) depends on experimentally controllable variables T and V/P
- an equilibrium state of the system
 - minimum of F or G rather than of U
 - equivalent to maximization of entropy of universe
- \blacksquare F or G decreases in the course of spontaneous process
 - holy grail of MD simulation

Microscopic definition of entropy

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W – number of ways ('microstates')in which a certain 'macrostate' may be realized(a certain occupation of energy levels of the system)
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microscopic entropy:

$$S = k_{\mathsf{B}} \cdot \ln W$$

 $(k_{\rm B} - {\rm universal \ Boltzmann \ constant})$

Microscopic definition of entropy



Microscopic definition of entropy

- S tells us something about the travel of the system through the configuration (phase) space
- \blacksquare popular idea S can be related to the order in the system
- low entropy few states are occupied
 - only a small part of configuration space accessible
 - ordered system
- high entropy many states are visited
 - extended regions of the configuration space covered
 - less ordered system

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Example – pile of books on a desk
Jan Černý (Charles University in Prague, Dept Cellular Biology):
anthropy – "entropy of human origin"
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Introduction

"... series of molecular structures generated by an MD simulation with the Berendsen thermostat does not represent the correct canonical ensemble" – what does this mean?

The phase space may be sampled (walked through) in various ways – just what is the correct way?

Closed system - canonical ensemble

system in thermal contact with the surroundings

- temperature rather than energy remains constant
- Boltzmann distribution of probability \mathcal{P}_i applies:

$$\mathcal{P}_{i} = \frac{\exp[-\beta \cdot \varepsilon_{i}]}{Q}$$

$$Q = \sum_{j} \exp[-\beta \cdot \varepsilon_{j}]$$

Q – canonical partition function (Zustandssumme)

derive the meaning of β – fall back to basic thermodynamics. . . :

$$\beta = \frac{1}{k_{\rm B}T}$$

Canonical partition function

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Q – seems to be purely abstract...
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but – to characterize the thermodynamics of a system Q is completely sufficient, all thermodynamic observables follow as functions of Q
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how to obtain Q?

- there will be 2 ways, depending on the system studied

partition function connects the microscopic and macroscopic world

Using the partition function

we can get all thermodynamic functions from Q:

$$U = \langle E \rangle = k_{\rm B} T^2 \frac{\partial \ln Q}{\partial T}$$

$$S = k_{\rm B} T \cdot \frac{\partial \ln Q}{\partial T} + k_{\rm B} \cdot \ln Q$$

$$F = -k_{\rm B} T \cdot \ln Q$$

$$P = k_{\rm B} T \cdot \left(\frac{\partial \ln Q}{\partial V}\right)_T \qquad \text{(equation of state)}$$

$$H = U + pV$$

$$G = F + pV = H - TS$$

Michal Otyepka (Palacký University Olomouc, Dept Physical Chemistry): "Just grab the partition function at the tail, and then you have everything!"

2 ways to thermodynamic properties

- simple molecules with 1 or few minima of energy
 - calculate the partition function (trans+vib+rot) and employ some approximations (IG+HO+RR)
 - derive properties from Q

Discrete systems

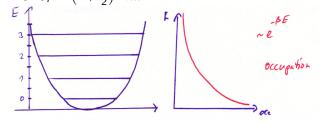
system with discrete energy levels ε_i – partition function:

$$Q = \sum_{i} \exp[-\beta \varepsilon_{i}]$$

Boltzmann distribution function: (prob. of system in state ε_i)

$$\mathcal{P}_i = \frac{1}{Q} \exp[-\beta \varepsilon_i]$$

example – HO: $\varepsilon_i = \left(i + \frac{1}{2}\right) \cdot \hbar \omega$



Using the partition function of 1 molecule

partition function of a large system - simplifications possible:

 \blacksquare system of *n* identical and indistinguishable particles (gas):

$$Q=\frac{q^n}{n!}$$

- necessary effort reduced greatly!
 - get the molecular partition function q (calculate 1 molecule, or 2)
 - obtain the ensemble partition function Q and all thermodynamic quantities

Simple molecules

(1 or few well characterized minima = dominant structures)

for a certain minimum - consider contributions to energy:

$$E = E^{\text{el}} + E^{\text{trans}} + E^{\text{rot}} + E^{\text{vib}}$$

partition function follows as

$$Q = \exp\left[-\beta \left(E^{\text{el}} + E^{\text{trans}} + E^{\text{rot}} + E^{\text{vib}}\right)\right] =$$

$$= \exp[-\beta E^{\text{el}}] \cdot \exp[-\beta E^{\text{trans}}] \cdot \exp[-\beta E^{\text{rot}}] \cdot \exp[-\beta E^{\text{vib}}] =$$

$$= Q^{\text{el}} \cdot Q^{\text{trans}} \cdot Q^{\text{rot}} \cdot Q^{\text{vib}}$$

or

$$\ln Q = \ln Q^{\mathsf{el}} + \ln Q^{\mathsf{trans}} + \ln Q^{\mathsf{rot}} + \ln Q^{\mathsf{vib}}$$

Electronic partition function

usually: quite high excitation energy

ightarrow electronic ground state only populated:

$$E^{el}(0) = 0$$
 arbitrarily

electronic partition function:

$$Q^{\rm el} = \exp[-\beta E^{\rm el}(0)] + \exp[-\beta E^{\rm el}(1)] + \dots \approx 1 + 0 + \dots = 1$$

so this may be neglected ©

Translational partition function

calculated for quantum-mechanical particle (mass m) in a 3D box: energy levels:

$$E_{n_x,n_y,n_z} = \frac{h^2}{8m} \left(\frac{n_x^2}{L_x^2} + \frac{n_y^2}{L_y^2} + \frac{n_z^2}{L_z^2} \right)$$

quantum numbers $n_i=1,2,\ldots$

partition function:

$$Q^{\mathsf{trans}} = \left(\frac{2\pi m k_{\mathsf{B}} T}{h^2}\right)^{\frac{3}{2}} \cdot V$$

Rotational partition function

calculated for a rigid rotor with moments of inertia I_x , I_y , I_z : energy levels:

$$E_J = B \cdot J(J+1)$$

quantum number $J=0,1,2,\ldots$, degeneracy of levels 2J+1 rotational constant $B_{\rm x}=rac{h^2}{8\pi^2I_{
m x}}$

$$Q^{\text{rot}} = \sum_{J=0}^{\infty} (2J+1) \exp \left[-J(J+1) \cdot \frac{B}{k_{\text{B}}T} \right]$$

for asymmetric top with rotational constants B_x , B_y , B_z :

$$Q^{\text{rot}} = \sqrt{\frac{\pi \left(k_{\text{B}}T\right)^3}{B_{\text{x}}B_{\text{y}}B_{\text{z}}}}$$

Vibrational partition function

calculated with harmonic vibrational frequencies ω_k of the molecule (computation of Hessian in the minimum of potential energy)

- each vibrational mode k is one harmonic oscillator

energy levels:
$$E_k^m = \left(m + \frac{1}{2}\right) \cdot \hbar \omega_k$$

where $E_k^0 = \frac{1}{2}\hbar\omega_k$ is zero point vibrational energy

partition function (using $\sum_{n=0}^{\infty} x^n = \frac{1}{1-x}$):

$$Q_k^{\text{vib}} = \sum_{m=0}^{\infty} \exp\left[-\beta \left(m + \frac{1}{2}\right) \hbar \omega_k\right] = \frac{\exp\left[-\frac{1}{2}\beta \hbar \omega_k\right]}{1 - \exp\left[-\beta \hbar \omega_k\right]}$$

each molecule: N-6 vibrational modes $\rightarrow N-6$ HOs example – H_2O : 3 modes (2 stretches, 1 bend)

Thermodynamic properties

for enthalpy -pV is needed - simple for IG:

$$pV = Nk_{\rm B}T$$

then, enthalpy and Gibbs free energy follow:

$$H = U + pV = U + Nk_BT$$

 $G = F + pV = F + Nk_BT$

thermal contributions – calculated by default
with many QCh and MD programs
whenever vibrational analysis is requested
reason – vibrational frequencies are computationally costly
while the thermodynamics is done 'for free'

Example – vibrational contributions

$$\ln Q_k = -rac{1}{2}eta\hbar\omega_k - \ln\left[1 - \exp[-eta\hbar\omega_k]
ight]$$
 $U_k = -rac{\partial \ln Q_k}{\partial eta} = \hbar\omega_k\left(rac{1}{2} + rac{1}{\exp[eta\hbar\omega_k] - 1}
ight)$

consider this for all of N-6 vibrational DOFs, dropping ZPVE:

$$U^{\mathsf{vib}} = \sum_{k=1}^{N-6} \left(\frac{\hbar \omega_k}{\exp[\beta \hbar \omega_k] - 1} \right)$$

$$F^{\text{vib}} = -k_{\text{B}} T \ln Q^{\text{vib}} = \sum_{k=1}^{N-6} k_{\text{B}} T \ln \left[1 - \exp[-\beta \hbar \omega_k] \right]$$

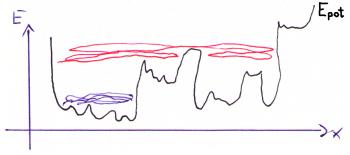
$$\frac{S^{\text{vib}}}{k_{\text{B}}} = \frac{U^{\text{vib}} - F^{\text{vib}}}{k_{\text{B}} T} = \sum_{k=1}^{N-6} \left(\frac{\beta \hbar \omega_k}{\exp[\beta \hbar \omega_k] - 1} - \ln \left[1 - \exp[-\beta \hbar \omega_k] \right] \right)$$

2 ways to thermodynamic properties

- simple molecules with 1 or few minima of energy
 - calculate the partition function (trans+vib+rot) (probably employ approximations IG+HO+RR)
 - derive properties from Q
- flexible molecules, complex molecular systems
 - quantum mechanical energy levels cannot be calculated
 - a single minimum of energy not meaningful
 - do MD (MC) simulation instead, to sample phase space
 - evaluate time averages of thermodynamic quantities

Systems with continuous distribution of energy

dynamics of molecules – at different tot. energies or temperatures, differently extended regions of conformational space are sampled



complex energy landscape $E_{\rm pot}(x)$ blue and red — trajectories at different total energies — different phase-space densities

Continuous systems – canonical ensemble

- every point in phase space a certain value of energy composed of $E_{pot} = E_{pot}(\vec{r})$ (force field), $E_{kin} = E_{kin}(\vec{p})$
- continuous energy levels infinitesimally narrow spacing
- canonical probability distribution function
 - probability to find the system in state with *E*:

$$\mathcal{P}(\vec{r}, \vec{p}) = \rho(\vec{r}, \vec{p}) = \frac{1}{Q} \cdot \exp\left[-\frac{E(\vec{r}, \vec{p})}{k_{\text{B}}T}\right]$$

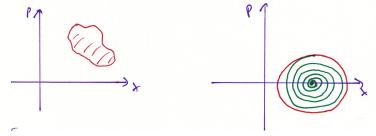
■ partition fction Q — integral over phase space rather than sum

$$Q = \int \exp\left[-\frac{E(\vec{r}, \vec{p})}{k_{\rm B}T}\right] \, \mathrm{d}\vec{r} \, \mathrm{d}\vec{p}$$

- perform MD simulation with a correct thermostat, or MC sim.
 - the goal is to obtain the right probability density ρ

Thermodynamic quantities - sampling

 $\rho(\vec{r}, \vec{p})$ – gives the probability of finding the system at (\vec{r}, \vec{p}) typically: system is sampling only a part of phase space $(\mathcal{P} \neq 0)$:



sampling in MD or MC

undamped and damped classical HO

Thermodynamic quantities - sampling

thermodynamic quantities – weighted averages:

$$\langle A \rangle = \frac{\int A \cdot \rho(\vec{r}, \vec{p}) \, d\vec{r} \, d\vec{p}}{\int \rho(\vec{r}, \vec{p}) \, d\vec{r} \, d\vec{p}}$$

why do MD? . . . obtain the correct phase-space density ho

– the density $\boldsymbol{\rho}$ is present in the the trajectory inherently

ightarrow e.g. structure with high ho occur more often

thermodynamic quantities in MD simulation – time averages:

$$\langle A \rangle = rac{1}{t_1 - t_0} \int_{t_0}^{t_1} A(t) \, \mathrm{d}t$$

- valid only if the simulation has sampled the canonical ensemble
 - → phase-space density is correct ergodic simulation

Thermodynamic quantities – sampling

Thus, we have the following to do:

- perform MD simulation (with correct thermostat!)
 - \rightarrow trajectory in phase space (simulation has 'taken care' of the phase-space density)
- get time averages of desired thermodynamic properties

$$\langle A \rangle = \frac{1}{t_1 - t_0} \int_{t_0}^{t_1} A(t) dt$$

Thermodynamic quantities – sampling

MD simulation serves to provide the phase-space density

$$\rho(\vec{r}, \vec{p}) = \frac{\exp[-\beta E(\vec{r}, \vec{p})]}{Q} \qquad \vec{r} = \{r_1, \dots, r_{3N}\}, \vec{p} = \{p_1, \dots, p_{3N}\}$$

which is the probability of system occurring at point (\vec{r}, \vec{p})

How long an MD simulation can we perform?

1 ps ightarrow 1,000 points in trajectory

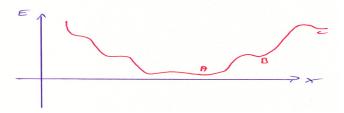
100 ns ightarrow 100M points – we cannot afford much more

let us think - we have 100M points

then, we have not sampled (\vec{r}, \vec{p}) for which $\rho(\vec{r}, \vec{p}) < 10^{-8}$

- ightarrow points with high energy will be never reached! (while low-energy region may be sampled well)
- actually, most of the points accumulate in low-energy regions, so this is much more serious!

Missing high-energy points in sampling



High-energy points B and C may be sampled badly

- a typical problem in MD simulations of limited length
- the corresponding large energies are missing in averaging
- when does this matter?
- no serious error for the internal energy exponential dependence of phase-space density kills the contribution

$$\rho(\vec{r}, \vec{p}) = \frac{\exp[-\beta E(\vec{r}, \vec{p})]}{Q}$$

Missing high-energy points in sampling

Free energies

- determine the spontaneity of process
- NVT canonical Helmholtz function F
- NPT 'canonical' Gibbs function G
- the relevant quantity always obtained depending on whether NVT or NPT simulation is performed
- much more pronounced sampling issues than e.g. for internal energy!

Missing high-energy points in sampling

$$F = -k_{\rm B}T \ln Q = k_{\rm B}T \ln \frac{1}{Q} =$$

$$= k_{\rm B}T \ln \frac{c^{-1} \cdot \iint \exp[\beta E(\vec{r}, \vec{p})] \cdot \exp[-\beta E(\vec{r}, \vec{p})] \, d\vec{r} \, d\vec{p}}{Q} =$$

$$= k_{\rm B}T \ln \frac{c^{-1} \cdot \iint \exp[\beta E(\vec{r}, \vec{p})] \cdot \exp[-\beta E(\vec{r}, \vec{p})] \, d\vec{r} \, d\vec{p}}{Q} =$$

$$= k_{\rm B}T \ln \iint \exp[\beta E(\vec{r}, \vec{p})] \cdot \rho(\vec{r}, \vec{p}) \, d\vec{r} \, d\vec{p} - \ln c =$$

$$= k_{\rm B}T \cdot \ln \left\langle \exp\left[\frac{E}{k_{\rm B}T}\right] \right\rangle - \ln c$$

serious issue – the large energy values enter an exponential, and so the high-energy regions may contribute significantly!

→ if these are undersampled, then free energies are wrong

- calculation of free energies impossible! special methods needed!