

# Molecular dynamics simulation

how to get things moving

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# Motivation

Consider a (bio)molecule in aqueous solution  
at ambient conditions

- structure is varying
- interactions are varying (H-bonds)
- the energy of the system is fluctuating
- description with a single, static structure – meaningless
- an interesting process may be going on 😊
- then, multiple 'structures' may be relevant

# State of the system

- (micro)state of a system:  
positions  $\vec{r}_i$  and momenta  $\vec{p}_i$  of all the atoms
- configuration space –  $3N$ -dimensional space of coordinates
- phase space –  $6N$ -dim. space of coords and momenta  $\{\vec{r}_i, \vec{p}_i\}$
- trajectory in phase space – sequence of points  $\{\vec{r}_i(t), \vec{p}_i(t)\}$   
passed by the system in course of time

# State of the system

example – 1D harmonic oscillator:

time course of coordinate and of velocity

$$r(t) = a \cdot \cos[\omega t]$$

$$v(t) = -a\omega \cdot \sin[\omega t]$$

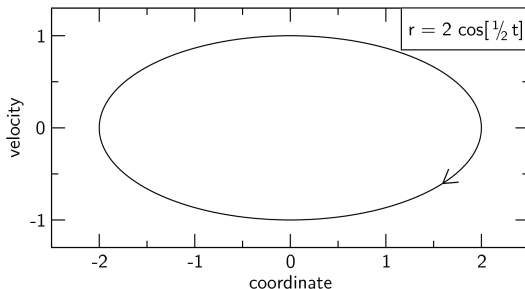
plot of velocity vs. coordinate – in 2D phase space:

elliptic trajectory

$$\left(\frac{x(t)}{a}\right)^2 + \left(\frac{v(t)}{a \cdot \omega}\right)^2 = 1$$

# State of the system

Example – 1D harmonic oscillator:

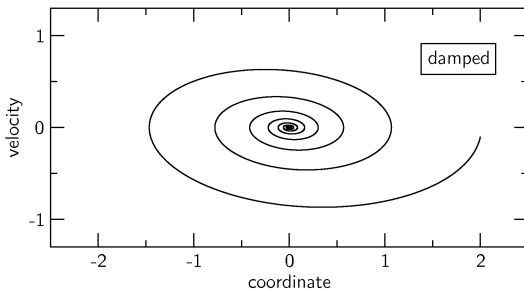


$$E_{\text{tot}} = E_{\text{kin}} + E_{\text{pot}} = \frac{1}{2} m \omega^2 a^2$$

**conservative system** – total energy remains conserved (constant)

# State of the system

Example – 1D harmonic oscillator:

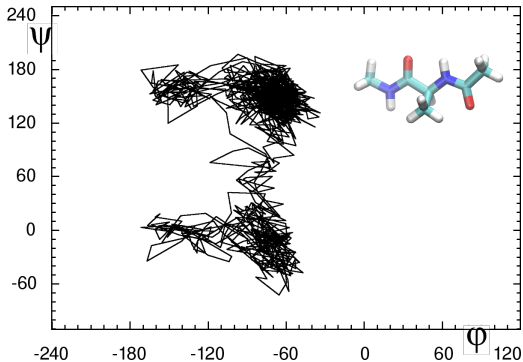


with **friction** or other **damping**

– the total energy of the system is decreasing

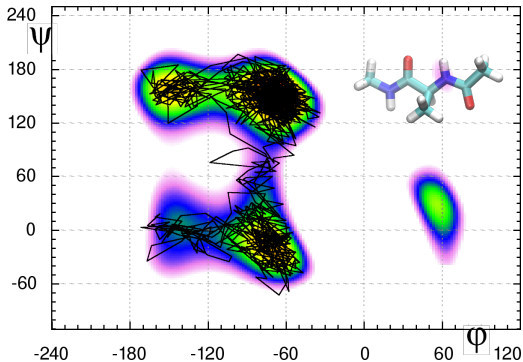
# State of the system

Example – alanine dipeptide in aqueous solution:  
config. space of dihedral angles  $\varphi - \psi$  (Ramachandran plot)



# State of the system

Example – alanine dipeptide in aqueous solution:  
config. space of dihedral angles  $\varphi - \psi$  (Ramachandran plot)





# Thermodynamic properties

Back to the **molecule in solution**:

MD simulation – we generate a trajectory in phase space

for some time  $\rightarrow$  **snapshots**  $\{\vec{r}_i(t_k), \vec{p}_i(t_k)\}$ ,

evaluate energy in time instants  $t_k$  ( $k = 1, \dots, M$ )  $\rightarrow E_k$

and calculate the average:

$$\langle E \rangle_t = \frac{1}{M} \sum_{k=1}^M E_k$$

generally – we obtain the average value of the property of interest,  
over all observed structures

to obtain the change of the property in a **reaction**:

we do this for the product and for the reactant,

obtain averages for both states and subtract them

# Thermodynamic properties

we obtain the trajectory by doing an MD simulation

– a good idea, but still there are issues:

- Do we have enough snapshots? – all relevant conformations?
- How do we consider experimental conditions – temperature?
- Suppose we know the structure of the reactant.  
How do we get the structure of the product?  
or even the whole reaction path?
- Does the average of energy provide useful information?  
What about free energies / entropy?

# Thermodynamic properties

Characteristics of (bio)molecular simulations:

- it is easy to derive the total energy – force field
- not so easy to make proper use of the energy function  
to get the **thermodynamic properties** right
- it is all about thermodynamics  
in possible contrast to quantum chemistry

# Thermodynamic properties

**time average** for energy and other properties of interest:

$$\langle A \rangle_t = \frac{1}{t_1 - t_0} \int_{t_0}^{t_1} A(t) dt$$

experimental sample – huge number of molecules,  
all relevant conformations of molecule/solvent are present  
– thermodynamic **ensemble**

How many molecules in the ensemble are found in  $\{\vec{r}_i, \vec{p}_i\}$ ?

→ **phase-space density** (per volume unit)  $\rho(\vec{r}, \vec{p})$

→ **ensemble average** can be calculated:

$$\langle A \rangle_e = \frac{\int A \cdot \rho(\vec{r}, \vec{p}) d\vec{r} d\vec{p}}{\int \rho(\vec{r}, \vec{p}) d\vec{r} d\vec{p}}$$

# Thermodynamic properties

experiment – ensemble average is always measured

simulation – a single molecule – time average available

simulation – system is considered **ergodic**

- passes through all points of phase space

constituting the real ensemble

provided the simulation is long enough

- implies:

$$\langle A \rangle_t = \langle A \rangle_e$$

the topic of **sampling**, danger of undersampling

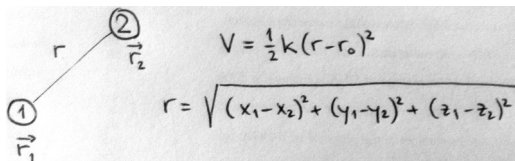
## Déjà vu – energy

$$\begin{aligned} E(R^N) = & \\ = & \frac{1}{2} \sum_i k_i (r_i - r_i^0)^2 + \frac{1}{2} \sum_j k_j^\vartheta (\vartheta_j - \vartheta_j^0)^2 + \frac{1}{2} \sum_n V_n \cdot \cos[n\omega - \gamma_n] \\ + & \sum_i^N \sum_{j=i+1}^N \left\{ 4\epsilon_{ij} \left( \left( \frac{\sigma_{ij}}{r_{ij}} \right)^{12} - \left( \frac{\sigma_{ij}}{r_{ij}} \right)^6 \right) + \frac{1}{4\pi\epsilon_0} \frac{q_i q_j}{r_{ij}} \right\} \end{aligned}$$

# Déjà vu – forces

$$V = V(r_{ij}) + V(r_{ik}) + V(r_{il}) + \dots$$

$$F_i^x = -\frac{\partial V}{\partial x_i} = -\frac{\partial V(r_{ij})}{\partial r_{ij}} \frac{\partial r_{ij}}{\partial x_i} - \frac{\partial V(r_{ik})}{\partial r_{ik}} \frac{\partial r_{ik}}{\partial x_i} - \frac{\partial V(r_{il})}{\partial r_{il}} \frac{\partial r_{il}}{\partial x_i} - \dots$$



$$\vec{F}_1 = -k(r_{12} - r_0) \cdot \frac{\vec{r}_{12}}{r_{12}}$$

# Equations of motion

total energy – **Hamilton function** (Hamiltonian):

$$H = T + V = \frac{1}{2} \frac{p^2}{m} + \frac{1}{2} k r^2$$

equations of motion in Hamilton's formalism:

$$\dot{r}_i = \frac{\partial H}{\partial p_i} \quad \dot{p}_i = -\frac{\partial H}{\partial r_i}$$

leading to **ordinary differential eqn** (ODE) of 2nd order

$$\dot{r} = \frac{\partial H}{\partial p} = \frac{p}{m} \rightarrow p = m\dot{r} \rightarrow \dot{p} = m \cdot \ddot{r}$$

$$\dot{p} = -\frac{\partial H}{\partial r} = -\frac{\partial V}{\partial r} = F$$

$$m \cdot \ddot{r} = F$$



# Equations of motion

example – harmonic oscillator:  $H = \frac{1}{2} \frac{p^2}{m} + \frac{1}{2} k r^2$

$$\dot{r} = \frac{\partial H}{\partial p} = \frac{p}{m}$$

$$\dot{p} = -\frac{\partial H}{\partial r} = F = -k \cdot r$$

equation of motion:

$$m \cdot \ddot{r} = -k \cdot r$$

- We will use the same concept – considering  $x, y, z$  of all of the atoms instead of  $r$ , and taking forces from the ‘long equation’
- Hamilton / Lagrange formalisms are more general
  - other coordinates than  $x, y, z$  of atoms may be used
  - internal coordinates ...

# Relevant differential equations

## 1st-order ODE

- generally:  $\dot{x} = f(x, t)$
- example:  $\dot{x} = -k \cdot x$
- solution:  $x(t) = A \cdot \exp[-k \cdot t]$
- e.g. radioactive decay, dynamics of populations

# Relevant differential equations

## 2nd-order ODE:

- $\ddot{x} = f(x, \dot{x}, t)$
- example: eqn of motion of harmonic oscillator  $\ddot{x} = -\frac{k}{m} \cdot x$
- with linear damping:  $\ddot{x} = -\zeta \cdot \dot{x} - \frac{k}{m} \cdot x$
- reduction of 2nd-order ODE to **two** 1st-order ODEs  
by introducing velocity  $v$ :

$$\begin{aligned}\dot{x} &= v \\ \dot{v} &= -\zeta \cdot v - \frac{k}{m} \cdot x\end{aligned}$$

these ODEs have to be solved numerically

# (Too) simple numerical solution

$$\ddot{r} = f(r, t)$$

common trick – Taylor expansion ( $\Delta t = t - t_0$ ):

$$r(t) = r(t_0) + \dot{r}(t_0) \cdot \Delta t + \frac{1}{2} \ddot{r}(t_0) \cdot \Delta t^2 + \dots$$

**Euler method** – 1st-order approximation:

$$r(t) \approx r(t_0) + \dot{r}(t_0) \cdot \Delta t$$

Numerical integration starts at time  $t_0$  – we make a **step  $\Delta t$** :

$$a(t_0) = -\frac{F}{m}$$

$$r(t_0 + \Delta t) = r(t_0) + v(t_0) \cdot \Delta t$$

$$v(t_0 + \Delta t) = v(t_0) + a(t_0) \cdot \Delta t$$

# Verlet – normal form

Euler method – too large numerical error  $\mathcal{O}(\Delta t^2)$

more accurate integration is needed

**Verlet method:**

Taylor expansion up to 2nd order,

derivation from two virtual steps, forwards and backwards:

$$r(t + \Delta t) = r(t) + \dot{r}(t) \cdot \Delta t + \frac{1}{2} \ddot{r}(t) \cdot \Delta t^2$$

$$r(t - \Delta t) = r(t) - \dot{r}(t) \cdot \Delta t + \frac{1}{2} \ddot{r}(t) \cdot \Delta t^2$$

add both equations – eliminate the velocity  $\dot{r}$ :

$$r(t + \Delta t) = 2 \cdot r(t) - r(t - \Delta t) + \ddot{r}(t) \Delta t^2$$

$$\ddot{r}(t) = a(t) = \frac{F(t)}{m} = -\frac{1}{m} \frac{\partial V}{\partial r}(t)$$

# Verlet – normal form

$$r(t + \Delta t) = 2 \cdot r(t) - r(t - \Delta t) + \ddot{r}(t)\Delta t^2$$

strange – not only  $r(t)$  and  $a(t)$  needed, but also  $r(t - \Delta t)$ ?

no problem – information equivalent to velocity, so that  
initial conditions may be converted:

$$r(t_0 - \Delta t) = r(t_0) - v(t_0) \cdot \Delta t$$

velocities – not in there explicitly, but may be obtained:

$$\dot{r}(t) = v(t) = \frac{r(t + \Delta t) - r(t - \Delta t)}{2 \cdot \Delta t}$$

(Verlet normal form)

## Verlet – normal form

program for 'astronomic' simulations: ( $\vec{F} = -1/r^2 \cdot \vec{r}/r$ )

```

/* initial "old" positions from initial velocities */
for (k=0; k<DIM; k++)
    r_old[k] = r[k] - v[k] * dt;

for (t=0.; t < CYCLES*PERIOD; t+=dt) {
    /* distance (from the Sun) */
    rnorm = sqrt(NORM2(r));
    /* gravitation force (on the comet)
    * f = -1 / r^2
    * multiply this by the unit vector in the direction of r
    * f = -1 / r^2 * vector(r) / r
    */
    for (k=0; k<DIM; k++)
        f[k] = - r[k] / CUB(rnorm);
    /* Verlet integrator */
    for (k=0; k<DIM; k++) {
        r_new = 2 * r[k] - r_old[k] + f[k] * SQR(dt);
        r_old[k] = r[k];
        r[k] = r_new;
    }
}

```

# Detailed balance

In equilibrium, this condition holds:

The rate of transitions from state  $i$  to state  $j$  is the same as from  $j$  to  $i$ , on average.

In the other words: the flux of probability from state  $i$  to state  $j$  is exactly balanced by the probability flux from  $j$  to  $i$ :

$$p_{i \rightarrow j} \cdot \rho_i^{\text{eq}} = p_{j \rightarrow i} \cdot \rho_j^{\text{eq}}$$

The **probability**  $\rho$  depends on the ensemble (which in turn depends on the conditions):

- isolated system – microcanonical ensemble:  
principle of equal a priori probabilities,  $\rho_i^{\text{eq}} = \rho$
- closed system – canonical ensemble:  $\rho_i^{\text{eq}} \propto \exp[-\beta E_i]$

Any MD algorithm/implementation shall observe detailed balance!



# Velocity Verlet

another, equivalent formulation

- positions calculated first using velocities

$$r(t + \Delta t) = r(t) + v(t) \cdot \Delta t + \frac{1}{2}a(t) \cdot \Delta t^2$$

- forces ( $\rightarrow$  accelerations) calculated in new positions, and new velocities obtained as

$$v(t + \Delta t) = v(t) + \frac{1}{2}(a(t) + a(t + \Delta t)) \cdot \Delta t$$

- next calculation of positions  $r \dots$

MD is started with the knowledge of  $r_0$  and  $v_0$

in every step,  $r(t + \Delta t)$  is calculated first

so that  $a(t + \Delta t)$  can be updated, to get  $v(t + \Delta t)$

# Velocity Verlet

VV – better numerical precision than normal Verlet  
numerical problem of normal Verlet

- adding a small but important term  $\ddot{r}(t_0)\Delta t^2$   
to a large term calculated as difference:  $2r(t) - r(t - \Delta t)$
- large relative uncertainty

desirable – use an algorithm that is mathematically equivalent but  
does not require to perform potentially problematic calculations

# Leap-frog

yet another equivalent formulation, similar to VV

–  $r$  and  $v$  are evaluated in an alternating fashion:

$$r(t), v(t + \frac{1}{2}\Delta t), r(t + \Delta t), v(t + \frac{3}{2}\Delta t), r(t + 2\Delta t) \dots$$

■ velocities at  $t + \frac{1}{2}\Delta t$  are obtained first:

$$v(t + \frac{1}{2}\Delta t) = v(t - \frac{1}{2}\Delta t) + a(t) \cdot \Delta t$$

■ then, positions are updated at  $t + \Delta t$ :

$$r(t + \Delta t) = r(t) + v(t + \frac{1}{2}\Delta t) \cdot \Delta t$$

So, accelerations have to be calculated at  $t, t + \Delta t, t + 2\Delta t \dots$

from forces, and positions are needed to compute forces

– in fact positions have to be known at the same  $t$  that we need  $a$

# Initial conditions

To start the MD

- the positions  $r_0$  and the velocities  $v_0$  have to be specified

First step – calculations of forces at  $r_0$  to get accelerations  $a_0$

Then – the integrator may provide  $r$  (and  $v$ ) at time  $t_0 + \Delta t$

To obtain a trajectory over a time interval  $T$ ,

we perform  $M$  steps

- we have to evaluate the forces on all atoms  $M = T/\Delta t$  times

Computational cost of the calculation of forces

determines how many steps we can afford to make

# $\Delta t$ – crucial parameter

Numerical issue:

- we neglect contributions in  $\Delta t^3$  and higher orders  $\rightarrow$  error per step in the order of  $\Delta t^3$  ( $\mathcal{O}(\Delta t^3)$ )
- keep the step short  $\rightarrow$  make the error small
- disadvantage: we may need too many steps to simulate certain time  $T$
- trade-off:  $\Delta t$  too long  $\rightarrow$  too large error  
dynamics may deviate, momentum may not be conserved. . .

# $\Delta t$ – crucial parameter

Chemical issue:

- fastest motion – hydrogen atoms, period around 10 fs
- rule of thumb – stable integration with  $\Delta t \leq$  fastest period (much more relaxed than in ‘astronomic’ simulations ☺)
- practically,  $\Delta t$  of 1 fs is used (2 fs with special treatment)

→ 1M calculations of forces needed for a trajectory of 1 ns  
large systems – multi-ns simulations routinely,  $\mu$ s possible

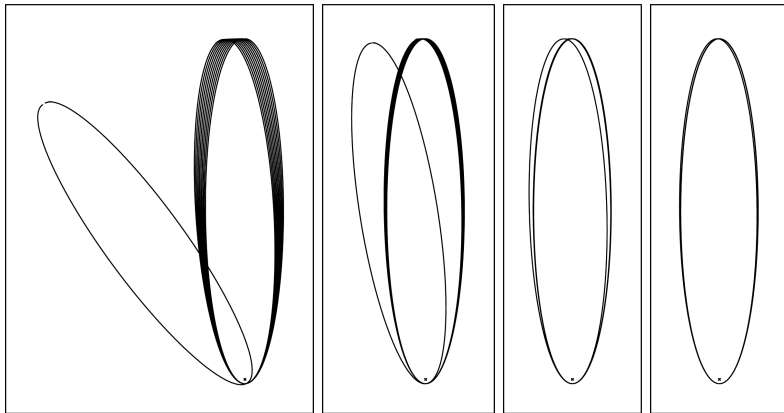
# $\Delta t$ – astronomic test

static heavy object (star) + moving light object (comet)  
+ gravity ( $F = -1/r^2$ ) → dynamics with Verlet integrator  
gravity – inverse-square law much like Coulomb between atoms  
exact trajectory – periodic along an ellipse with star in 1 focus

simulation – four different values of the time step:  
 $1 \times 10^{-5}$ ,  $2 \times 10^{-5}$ ,  $5 \times 10^{-5}$  and  $10 \times 10^{-5}$  of the orbital period  
→ 100k, 50k, 20k and 10k steps per period – mmore than in MD

# $\Delta t$ – astronomic test

first 10 orbits are shown, and then the 100th shown again



credit for the idea: Jiří Kolafa, Prague



## $\Delta t$ – astronomic test

shortest step – reasonable trajectory, small deviation

2nd-longest step – error becomes evident

longest step – large deviation

**important** – trajectory is precessing (ellipse is ‘rotating’),

but it remains elliptic whatever the time step is

constant remain also **total energy** and **orbital period**

– consequence of **reversibility** of Verlet:

if we reverse the course of time ( $\Delta t \rightarrow -\Delta t$ ),

we will simulate towards initial conditions of the trajectory

generally – energy in Verlet fluctuates (with longer step),

but it does not **drift**

# Verlet or something better?

Verlet – very approximative yet still routinely used for MD

why? – because it is efficient – why?

- forces on atoms ( $\rightarrow$  accelerations) calculated only  $1\times$  per step
- no higher derivatives of positions are involved

more accurate methods to integrate ODEs are available, and are used in some applications, if improved accuracy is required

straightforward – involve extra terms from Taylor expansion  
– hardly ever done, there are other ways to improve accuracy...

## Gear: predictor–corrector

- provides solution correct to an order of choice
- new positions are calculated (**predicted**) from Taylor expansion using a certain number of previous steps
- then, forces ( $\rightarrow$  accelerations) are calculated in the predicted positions
- accelerations used to make **correction** of positions
- additional computational effort, decreased efficiency
- accuracy may be improved significantly, longer step possible
- still, only 1 calculation of forces per step

# Gear: predictor–corrector

$n$ th-order Gear integrator:

coords of all atoms  $\vec{r}$  and their derivatives up to the order of  $n - 1$ :

$$R = \begin{pmatrix} \vec{r} \\ \dot{\vec{r}} \cdot \Delta t \\ \ddot{\vec{r}} \cdot \frac{1}{2} \Delta t^2 \\ \ddot{\vec{r}} \cdot \frac{1}{6} \Delta t^3 \end{pmatrix}$$

for the 4th-order method

initialization:  $\vec{r}$  and  $\dot{\vec{r}}$  from init. conditions,  $\ddot{\vec{r}}$  calculated from forces

– 1 calculation of forces required at start

higher derivatives may be set to zero

## Gear: 1: prediction

MD step at time  $t$  starts with **prediction**  
of coordinates+derivatives at time  $t + \Delta t$ :

$$R_p(t + \Delta t) = \begin{pmatrix} 1 & 1 & 1 & 1 \\ 0 & 1 & 2 & 3 \\ 0 & 0 & 1 & 3 \\ 0 & 0 & 0 & 1 \end{pmatrix} \cdot R(t)$$

- the matrix contains binomial coefficients
- the calculation passes a polynomial of order  $n - 1$   
through the previous  $n$  points of the trajectory  
(at  $t, t - \Delta t, \dots, t - (n - 1)\Delta t$ )  
and generates a point on this polynomial after  $\Delta t$
- prediction may be good for continuous force functions
- no calculation of force up to this point!

## Gear: 2: error

next, we calculate the **error** of the prediction:

we obtain the force at the predicted position,

and compare it with the force predicted in step 1  $\rightarrow$  error

$$E = \frac{1}{2} \left( \frac{\vec{f}(\vec{r}_p)}{m} - \ddot{\vec{r}}_p \right) \Delta t^2$$

$E$  – vector with as many components as the vector of coordinates  
(every coordinate with its derivatives has ‘its own’ error)

## Gear: 3: correction

Finally, using the error  $E$ , we calculate the **corrected** coordinates and derivatives as

$$R(t + \Delta t) = R_p(t + \Delta t) + E \cdot \begin{pmatrix} a_0 \\ a_1 \\ a_2 \\ a_3 \end{pmatrix}$$

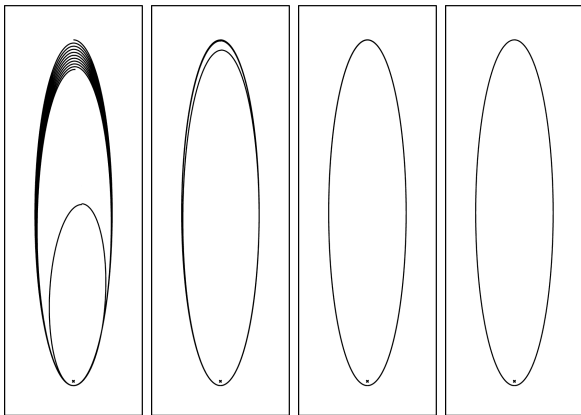
coefficients  $a_0, a_1 \dots a_{n-1}$

- estimated to prevent the accumulation of integration errors
- may be looked up in tables
- for 4th-order method for 2nd-order ODE  
and forces not depending on velocities:

$$a_0 = \frac{1}{6}, a_1 = \frac{5}{6}, a_2 = 1 \text{ and } a_3 = \frac{1}{3}$$

## Gear – astronomic test

$\Delta t = 10 \times 10^{-5}$ ,  $5 \times 10^{-5}$ ,  $2 \times 10^{-5}$  and  $1 \times 10^{-5}$  of correct period

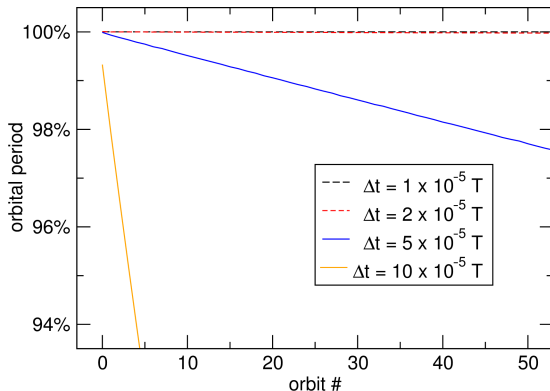




## Gear – astronomic test

- Gear may provide more accurate trajectories than Verlet, with the same  $\Delta t$
- perfect trajectories with the two shortest time steps (Verlet showed deviations even with the shortest step)
- incorrect behavior with the second-largest step, just wrong with the longest one
- different **character of deviation** than with Verlet: the elliptic trajectory of the comet is getting 'shorter', rather than precessing
- important: the orbital **period** is becoming shorter, and total **energy is decreasing**

# Gear – astronomic test



general observation:

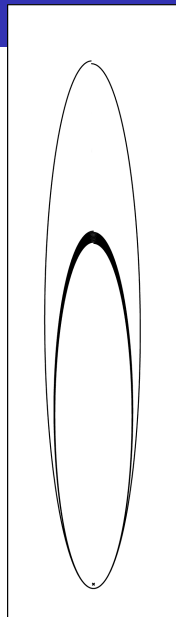
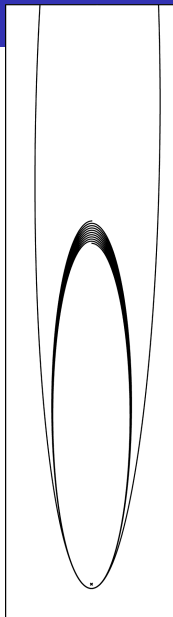
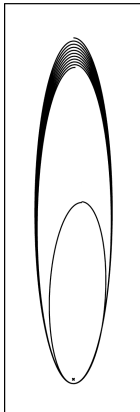
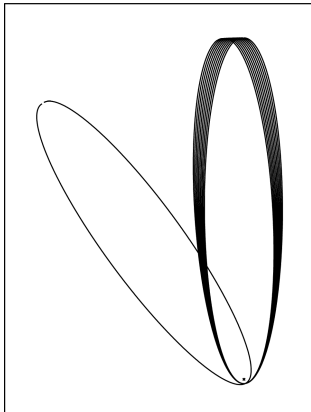
energy will decrease or increase (**drift**) in the simulation

this may be negligible with longer step / higher-order Gear

Gear: **not** reversible, does **not** conserve energy.

# Gear – higher order?

Verlet and Gear 4th, 5th and 6th order ( $\Delta t = 10 \times 10^{-5} T$ )



## Gear – higher order?

note: Verlet corresponds to 3th-order Gear formally  
higher derivatives in calculation improve the results only slightly  
– the drift of energy is slower but still unsatisfactory

general observations:

- when making the step shorter,  
results of higher-order methods will improve faster
- when making the time step longer,  
higher-order methods are more prone to fail completely  
while lower-order methods are more robust
- higher-order integrators are a good choice  
if accurate trajectories are desired
- lower-order or Verlet integration is sufficient for applications  
with 'weaker' requirements – typically, MD

# Runge–Kutta integration

Runge–Kutta methods – numerical integrators of 1st-order ODEs  
classical 4th-order method **RK4**:

- 4 calculations of the derivative in every step
- points at which the derivative is calculated
  - chosen depending on the previous calculations,
  - the first is done at the start of the integration step

$$g_0 = \dot{r}(r(t))$$

$$g_1 = \dot{r}\left(r(t) + \frac{1}{2}g_0\Delta t\right)$$

$$g_2 = \dot{r}\left(r(t) + \frac{1}{2}g_1\Delta t\right)$$

$$g_3 = \dot{r}\left(r(t) + g_2\Delta t\right)$$

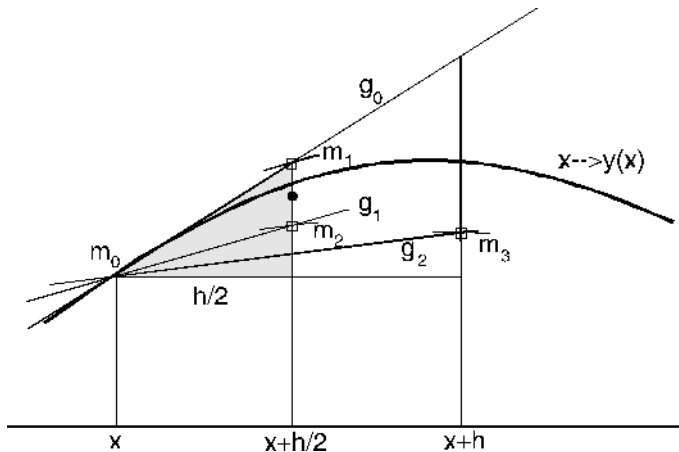
We calculate the value of the function at time  $t + \Delta t$   
using a weighted average of the obtained derivative values:

$$r_{n+1} = r_n + \frac{1}{6}(g_0 + 2g_1 + 2g_2 + g_3) \cdot \Delta t$$

# the RK4 method

derivative of  $y$  is calculated at points  $m_0$ ,  $m_1$ ,  $m_2$  and  $m_3$

calculated derivatives  $g_0, \dots, g_3$  are shown as arrows



# the RK4 method

- actually predictor–corrector with 4 predictions per step
- 4 calculations of the derivative needed per step
- error per step reduced to  $\mathcal{O}(\Delta t^5)$  – 4th-order method
- solves 1st-order ODEs much like the Euler method does
- to solve Newton eqns of motion (2nd-order ODEs)
  - eqns are converted to system of two 1st-order eqns
  - positions and velocities of atoms are propagated:

$$\begin{aligned}\dot{\vec{r}} &= \vec{v} \\ \dot{\vec{v}} &= \frac{\vec{f}}{m}\end{aligned}$$

# the RK4 method

Can RK4 be used directly somewhere in computational chemistry?

Yes! Let us propagate a time-dependent Schrödinger equation

– 1st-order ODE for the wave function  $\Psi$  of the system:

$$\frac{\partial \Psi}{\partial t} = -\frac{i}{\hbar} \hat{H} \Psi$$

- we express  $\Psi$  as linear combination of suitable basis functions:

$$\Psi = \sum_m c_m \varphi_m$$

- Hamiltonian is a matrix of elements between basis functions:

$$H_{mn} = \langle \varphi_m | \hat{H} | \varphi_n \rangle$$

- we calculate the derivative with matrix multiplication as  $H \cdot \Psi$